

# Combining laser light with synchrotron radiation

# Properties of Laser Light



coherence

interferometer

measuring distances

measuring wavelength

speckle measurements

structure reconstruction

# Properties of Laser Light



coherence



intensity (brightness)

melting of surfaces

study of phase transitions (pump probe)

plasma production

properties of the plasma state

ion and cluster production (laser ablation)

electron production (electron gun for linac)

higher harmonic generation

strong field effect

# Properties of Laser Light

- coherence
- intensity (brightness)
- spectral resolution (monochromatic)

high resolution absorption spectra

defined excitation of arbitrary lines

line width of continuous lasers  $< 1\text{MHz}$  ( $4\text{neV}$ )

example Eu

hyper fine structure (quadrupole)  $192\text{ MHz}$  ( $0.8\mu\text{eV}$ )

isotope shift  $2.8\text{ GHz}$  ( $12\mu\text{eV}$ )

# Properties of Laser Light

- ☉ coherence
- ☉ intensity (brightness)
- ☉ spectral resolution (monochromatic)
- ☉ spacial resolution (collimation)

high spacial coherence

→ point like source

→ optimal spot area possible  $\sim \lambda^2$

# Properties of Laser Light

- coherence
- intensity (brightness)
- spectral resolution (monochromatic)
- spacial resolution (collimation)
- temporal resolution

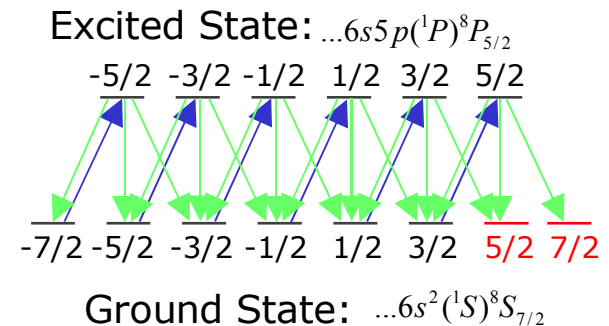
observing fast processes

ns:	atomic lifetimes	quality switched laser
ps:	melting processes	mode locked laser
fs:	chemical reactions	pulse compression
as:	photon absorption	high harmonic generation

# Properties of Laser Light

- ☪ coherence
- ☪ intensity (brightness)
- ☪ spectral resolution (monochromaticity)
- ☪ spacial resolution (collimation)
- ☪ temporal resolution
- ☪ polarization

populate magnetically aligned states



magneto optical trap (MOT)

# Properties of Laser Light

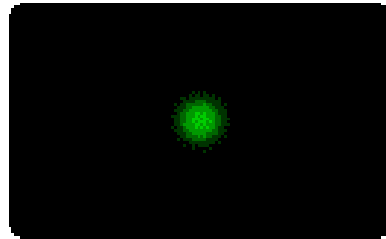
- coherence
- intensity (brightness)
- spectral resolution (monochromaticity)
- spacial resolution (collimation)
- temporal resolution
- polarization

## How to Produce Laser Light



# Laser Principle

Light **A**mplification by  
**S**timulated **E**mission of **R**adiation



Stimulated emission amplifies the  
existing light wave

# What is light really?

A stream of particles?

A wave?

Something else?

# Maxwell's equations

Gauß' Law

$$\nabla \cdot \vec{D} = \rho$$

No magnetic monopoles

$$\nabla \cdot \vec{B} = 0$$

Faraday's law

$$\nabla \times \vec{E} = -\frac{\delta \vec{B}}{\delta t}$$

Ampère's Law

$$\nabla \times \vec{H} = \vec{J} + \frac{\delta \vec{D}}{\delta t}$$

$$\vec{D} = \epsilon_0 \vec{E} + \vec{P}$$

$$\vec{B} = \mu_0 (\vec{H} + \vec{M})$$

# Wave equations

$$\nabla^2 \vec{E} = \epsilon \mu \frac{\delta^2 \vec{E}}{\delta t^2} \quad \nabla^2 \vec{H} = \epsilon \mu \frac{\delta^2 \vec{H}}{\delta t^2}$$

Free Wave:

$$\vec{E} = \sum_i \vec{A}_i e^{i(\omega_i t + \vec{k}_i \cdot \vec{r})} + c.c.$$

Speed of light:

$$v = \frac{\omega_i}{k_i} = \frac{1}{\sqrt{\epsilon \mu}}$$

Transversal:

$$\vec{E} \cdot \vec{k} = 0$$

# What is light really?

A stream of particles?

A wave?

Something else?

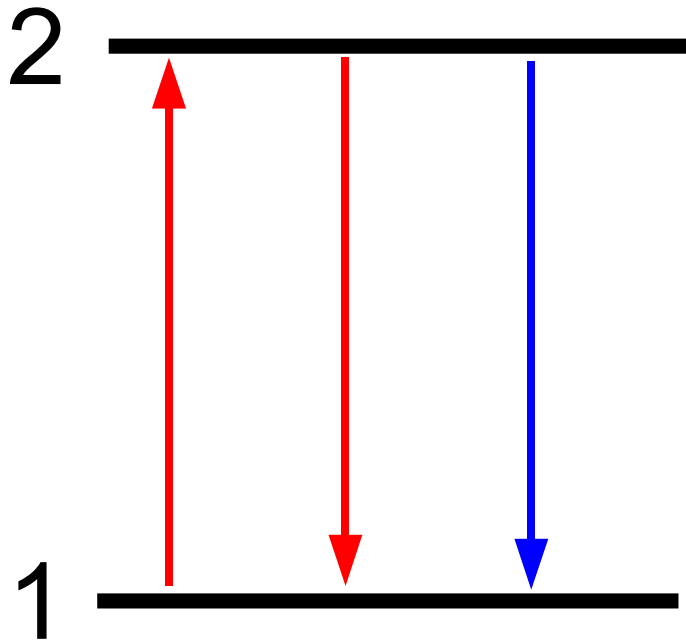
# Semiclassical approximation

In most cases, the interaction of light with matter can be modeled by treating the **light as a classical wave** and the **atoms quantummechanically**

Light: The number of photons per mode is usually high

Matter: At room temperature most atoms are in the lowest states

# Transition probabilities



Absorption  $W_{12}$

Stimulated  
emission  $W_{21}$

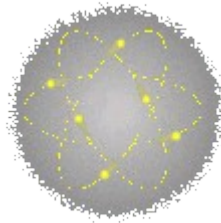
Spontaneous  
emission  $A$

# Calculating the transitions

Green light wavelength  $5000\text{\AA}$



$\sim 1\text{\AA}$





# Approximations

- Wavelength much greater than atom  
(Dipole approximation)
- Interaction over many periods  
(time average)
- Weak perturbation  
(first order perturbation theory)

# Atomic wave functions

Two energy eigenstates of an atomic system:

$$\Psi_1(\vec{r}, t) = u_1(\vec{r}) \exp\left[-i(E_1/\hbar)t\right]$$

$$\Psi_2(\vec{r}, t) = u_2(\vec{r}) \exp\left[-i(E_2/\hbar)t\right]$$

Superposition of eigenstates is a solution of the time dependent Schrödinger-equation:

$$\Psi = a_1 \Psi_1 + a_2 \Psi_2$$

with  $|a_1|^2 + |a_2|^2 = 1$

# Perturbation theory

Time independent atomic Hamiltonian:  $H_0$

Time dependent perturbation:

$$H'(t) = \vec{\mu} \cdot \vec{E} = e \vec{r} \cdot \vec{E}_0 \sin \omega t$$

Ansatz:

$$\Psi = a_1(t) \Psi_1 + a_2(t) \Psi_2$$

Gives in first order:

$$\frac{d}{dt} a_2 = (1 / i \hbar) H_{12} \exp(i \omega_0 t)$$

$$\text{with: } H_{12}(t) = \int u_1^* H' u_2 \quad \text{and} \quad \omega_0 = (E_2 - E_1) / \hbar$$

# Absorption probability

Integration over a long time gives the **transition probability**:

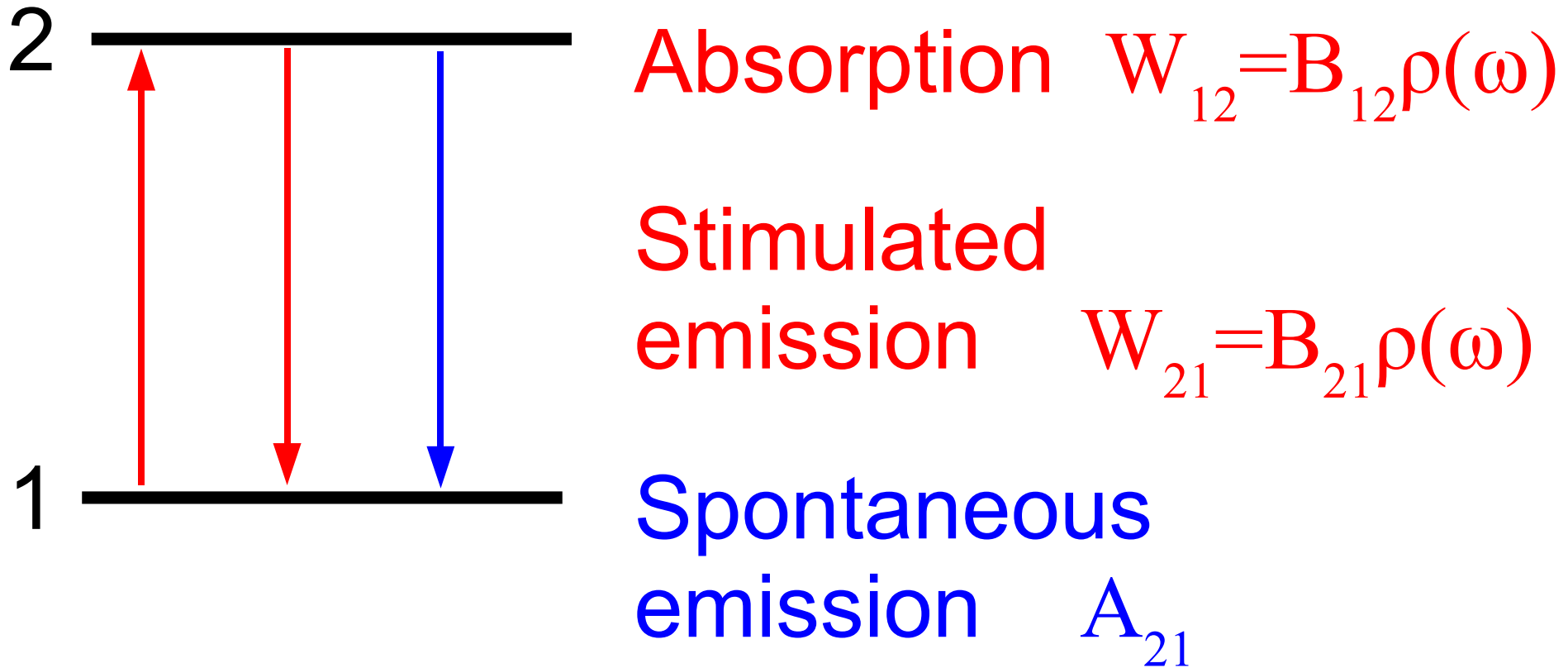
$$W_{12} = \frac{|a_2(t)|^2}{t} = \frac{\pi^2}{3\hbar^2} E_0^2 |\mu_{21}|^2 \delta(\omega - \omega_0)$$

with:

$$\vec{\mu}_{12} = \int u_2^* e \vec{r} u_1 dV$$

$W_{12}$  depends on light intensity and atomic properties

# Einstein Parameters



$\rho(\omega)$ : Spectral energy density

# Spontaneous Emission $A_{21}$

Spontaneous emission cannot be calculated in the semiclassical approximation.

→ QED

# Einstein's argumentation

Thermal equilibrium:

$$A_{21} N_2 + B_{21} \rho(\nu) N_2 = B_{12} N_1 \rho(\nu)$$

Boltzmann distribution:

$$N_i = (g_i N / Z) e^{-E/kT}$$

Blackbody radiation:

$$\rho(\nu) = \frac{8\pi\nu^2}{c^3} \frac{h\nu}{e^{h\nu/kT} - 1}$$

# Einstein parameters

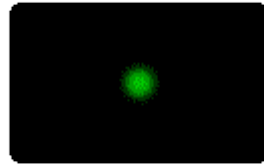
$$B_{12} = \frac{g_{12}}{g_{21}} B_{21}$$

$$A_{21} = \frac{8 \pi h \nu^3}{c^3} B_{21}$$



# Laser Principle

## Light **A**mplification by **S**timulated **E**mission of **R**adiation

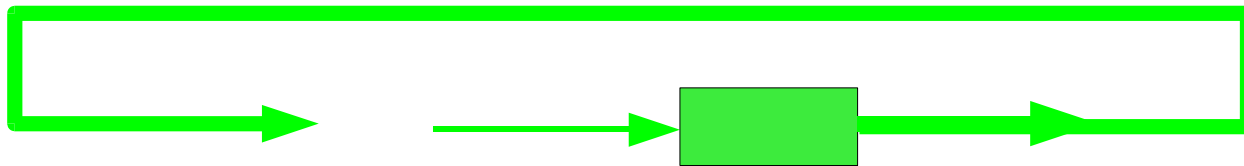


- Stimulated emission amplifies the existing light wave
- often within an optical resonator
- active medium excited by:
  - flash lamp
  - discharge
  - laser

# Resonator

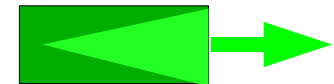
Laser with optical resonator:

- feedback turns an Amplifier into an Oscillator

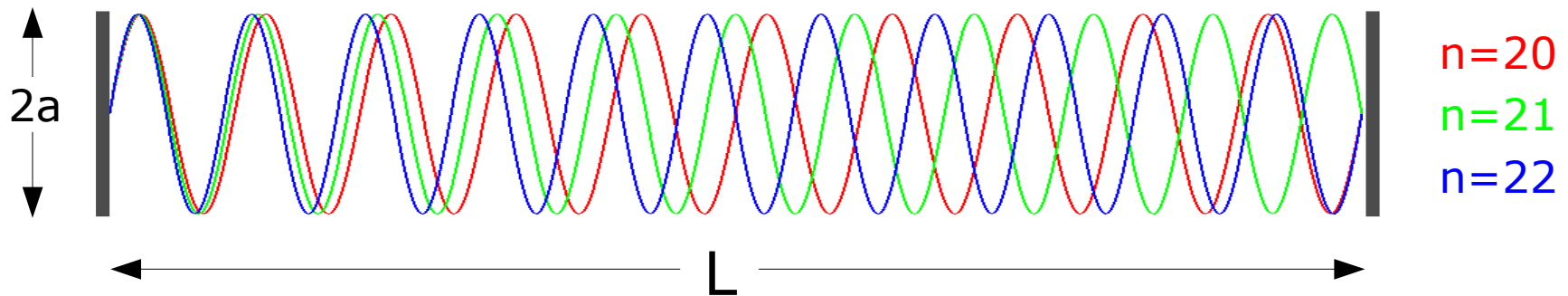


Laser without optical resonator:

- amplified spontaneous emission (S)ASE
- amplification of a random mode
- amplification of a seeding beam

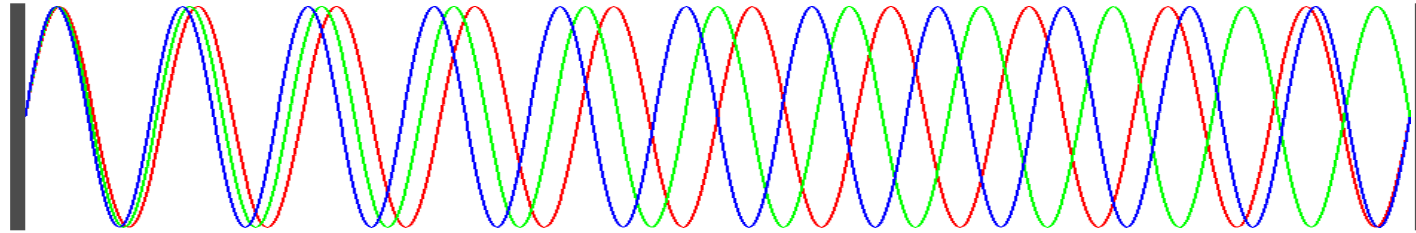


# Fabry-Perot Resonator



- standing waves between flat mirrors
- integer of half a wavelength must fit into the resonator  $\lambda = \frac{2L}{n}$
- $L=50\text{cm}$ ,  $\lambda=500\text{nm}$ :  $n=20\ 000$
- hard to align long resonators  $F = \frac{a^2}{\lambda L}$
- high diffraction losses for small Fresnel-number

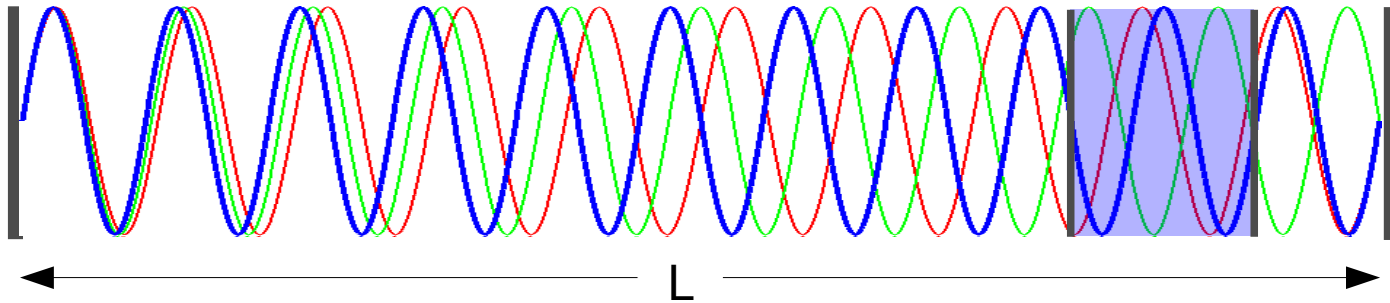
# Using the Modes



What do we need?

- spectral resolution
- coherence
  - Select a single mode
- high intensity
- short pulses
  - Couple the modes to a short pulse

# Mode Selection



mode separation (free spectral range):

$$\bullet \quad \Delta \nu = \frac{c}{2L}$$

• 300 Mhz for  $\frac{1}{2}$  m cavity

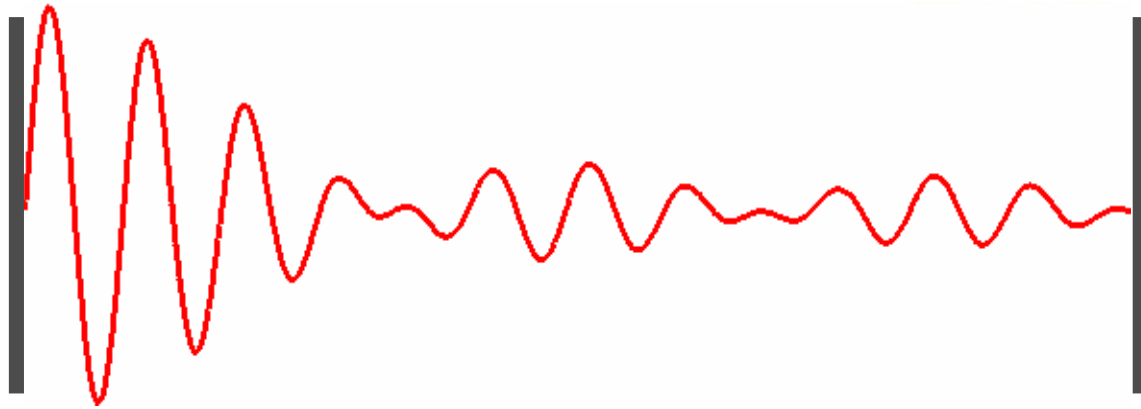
• 100 GHz for the thin etalon in the Ti:Sa

$\sim 1\text{nm}$  at  $\lambda = 600\text{nm}$

# Mode Coupling

producing pulse trains by fixing  
the phases of the modes

coupling 6 waves  $n=20-25$



coupling 101 waves  $n=20000-20100$

